Karlsruhe Institute of Technology, 16 January '12

Particle Physics and the first LHC results

> Guido Altarelli Roma Tre/CERN

In honour of Julius Wess





## Julius Wess

A great man A superb scientist A friend





# The first great result is that the LHC has worked very well in 2011!



This is great news for particle physics !!

The Standard Model cannot be the whole story

The SM is a low energy effective theory (nobody can believe it is the ultimate theory)

It happens to be renormalizable and highly predictive. And is (too) well supported by the data.

But even just as a low energy effective theory the SM is not satisfactory

In fact it is not completely verified: its simplest Higgs sector is so far only a conjecture and is problematic

and we expect New Physics at higher energies

not only from the GUT or Planck scales but also from the TeV scale (LHC!) hierarchy, dark matter...



## The Higgs problem is central in particle physics today

The main problems of the SM show up in the Higgs sector

$$V_{Higgs} = V_0 - \mu^2 \phi^{\dagger} \phi + \lambda (\phi^{\dagger} \phi)^2 + [\overline{\psi}_{Li} Y_{ij} \psi_{Rj} \phi + h.c.]$$
  
Vacuum energy
  
V\_{0exp}~(2.10^{-3} \text{ eV})^4
  
Possible instability
  
depending on m\_H
  
Origin of evendentia

Origin of quadratic divergences. Hierarchy problem The flavour problem: large unexplained ratios of Y<sub>ij</sub> Yukawa constants That some sort of spontaneous symmetry breaking mechanism is at work has already been established (couplings symmetric, spectrum totally non symmetric) The question is on the nature of the Higgs mechanism/particle(s)

- One doublet, more doublets, additional singlets?
- SM Higgs or SUSY Higgses
- Fundamental or composite (of fermions, of WW....)
- Pseudo-Goldstone boson of an enlarged symmetry
- A manifestation of extra dimensions (fifth comp. of a gauge boson, an effect of orbifolding or of boundary conditions....)
- Some combination of the above

## Alternative forms of EW symmetry breaking A vast literature

## Examples:

- SUSY Higgs
- Little Higgs
- Higgs from Extra Dim's
- Higgsless models
- Composite Higgs
- • •

Crosstalk with string theory:

Except for SUSY, common ingredients are: the Higgs is a pseudo Goldstone boson of an enlarged symmetry ---> new vector bosons Z', W', ρ'... Non perturbative sectors limit predictivity and all need an UV completion



#### Can we do without the Higgs?

Suppose we take the gauge symmetric part of the SM and put masses by hand.

Gauge invariance is broken explicitly. The theory is no more renormalizable. One loses understanding of the observed accurate validity of gauge predictions for couplings.

Still, what is the fatal problem at the LHC scale?

The most immediate disease that needs a solution is the occurrence of unitarity violations in some amplitudes



With no Higgs unitarity violations for  $E_{CM} \sim 1-3$  TeV

Unitarity implies that scattering amplitudes cannot grow indefinitely with the centre-of-mass energy s

In the SM, the Higgs particle is essential in ensuring that the scattering amplitudes with longitudinal weak bosons (W<sub>L</sub>, Z<sub>L</sub>) satisfy (tree-level) unitarity constraints [Veltman, 1977; Lee-Quigg-Thacker, 1977; ...] Zwirner

An example:  $\mathcal{A}(W_L^+ W_L^- \to Z_L Z_L) \quad (s \gg m_W^2)$ 



#### Can we do without the Higgs?

Suppose we take the gauge symmetric part of the SM and put masses by hand.

Gauge invariance is broken explicitly. The theory is no more renormalizable. One loses understanding of the observed accurate validity of gauge predictions for couplings.

Still, what is the fatal problem at the LHC scale?

The most immediate disease that needs a solution is the occurrence of unitarity violations in some amplitudes

To avoid this either there is one or more Higgs particles or some new states (e.g. new vector bosons)

Thus something must happen at the few TeV scale!!

A crucial question for the LHC

What saves unitarity?

• the Higgs

 some new vector boson W', Z' KK recurrences resonances from a strong sector

•••••



The main LHC results so far

• A robust exclusion interval for the SM Higgs. Only a narrow window is left below 600 GeV: 115.5-127 GeV.

ATLAS, CMS

Plus some indication for  $m_H \sim 125 \text{ GeV}$ 

- No evidence of new physics, althouh a big chunk of new territory has been explored
- Important results on B and D decays from LHCb [e.g.  $B_s$ ->J/ $\Psi \phi$ ,  $B_s$ -> $\mu \mu$ , .... CP viol in D decay]
- Two heavy ion runs so far (ALICE)
- Forward pp physics (TOTEM)

#### The 95% exclusion intervals for the light Higgs





Some "excess" was reported in the allowed  $m_H$  window

Is this the Higgs signal?

We hope yes, but the present evidence could still evaporate with more statistics

We need to wait for the 2012 run





Also in CMS there is an excess, but smaller (2.6  $\sigma$ )



 $\oplus$ 

#### Here is an attempt to put all the evidence together



 $\oplus$ 

#### Do the masses really coincide?



#### all data except CMS

#### all data except ATLAS



#### Peaks come and go! CMS History: $H \rightarrow \gamma \gamma$ <sup>Paus</sup> <sup>Zuric</sup>

#### Paus Zurich Jan. '12



#### A moderate enhancement of the $\gamma\gamma$ rate may be indicated





The SM Higgs is close to be observed or excluded!

Either the SM Higgs is very light (115.5 - 127 GeV) or rather heavy (i.e. > 600 GeV)

The range  $m_H = 115.5 - 127$  GeV is in agreement with precision tests, compatible with the SM and also with the SUSY extensions of the SM

 $m_{\rm H} \sim 125$  GeV is what you expect from a direct interpretation of EW precision tests: no fancy conspiracy with new physics to fake a light Higgs while the real one is heavy

 $m_H > 600$  GeV would point to the conspiracy alternative

## Theoretical bounds on the SM Higgs mass



If the SM would be valid up to  $M_{GUT}$ ,  $M_{Pl}$  with a stable vacuum then  $m_{H}$  would be limited in a small range depends on  $m_t$  and  $\alpha_s \longrightarrow 130$  GeV  $< m_H < 180$  GeV

But metastability (with sufficiently long lifetime) is enough!



In the absence of new physics, for  $m_H \sim 125$  GeV, the Universe becomes metastable at a scale  $\Lambda \sim 10^{10}$  GeV GeV But the SM remains viable up to  $M_{Pl}$  (Early universe implications)



Note that  $\lambda$ =0 at the Planck scale (and no physics in between) implies m<sub>H</sub> ~ 130 GeV depending on m<sub>t</sub> and  $\alpha_s$ 

$$m_h > 130 \,\text{GeV} + 1.8 \,\text{GeV} \left(\frac{m_t - 173.2 \,\text{GeV}}{0.9 \,\text{GeV}}\right) - 0.5 \,\text{GeV} \left(\frac{\alpha_s(M_Z) - 0.1184}{0.0007}\right) \ \pm 3 \ \text{GeV}$$

not far from 125 GeV Elias-Miro' et al, Holthausen et al, Wetterich '11

## The Standard Model works very well

So, why not find the Higgs and declare particle physics solved? Why one expects New Physics?

Because of both:

**Conceptual problems** 

- Quantum gravity
- The hierarchy problem
- The flavour puzzle

#### ....

and experimental clues:

- Neutrino masses
- Coupling unification
- Dark matter
- Baryogenesis
- Vacuum energy
- some experimental anomalies: (g-2), hints

Some of these problems point at new physics at the weak scale: eg Hierarchy Dark matter (perhaps)

> insert here your preferred hints



Neutrino masses are really special!

 $h_{\rm t}/(\Delta m_{\rm atm}^2)^{1/2} \sim 10^{12}$ 

Massless v's?

- no  $v_R$
- L conserved

## Small v masses?

- $v_R$  very heavy
- L not conserved

Very likely: v's are special as they are Majorana fermions

## A very natural and appealing explanation:

v's are nearly massless because they are Majorana particles and get masses through L non conserving interactions suppressed by a large scale M ~  $M_{GUT}$ 

 $m_v \sim \frac{m^2}{M}$  m: ≤  $m_t \sim v \sim 200$  GeV M: scale of L non cons.

Note:

$$m_v \sim (\Delta m_{atm}^2)^{1/2} \sim 0.05 \text{ eV}$$
  
m ~ v ~ 200 GeV



M ~ 10<sup>14</sup> - 10<sup>15</sup> GeV

## Neutrino masses are a probe of physics at M<sub>GUT</sub> !

How to prove that v's are Majorana fermions?

All we know from experiment on  $\nu$  masses strongly indicates that  $\nu$ 's are Majorana particles and that L is not conserved (but a direct proof still does not exist).



Detection of  $0\nu\beta\beta$  (neutrinoless double beta decay) would be a proof of L non conservation ( $\Delta L=2$ ). Thus a big effort is devoted to improving present limits and possibly to find a signal.

Heidelberg-Moscow, Cuoricino-Cuore, GERDA, •••••

Baryogenesis by decay of heavy Majorana v's BG via Leptogenesis near the GUT scale  $T \sim 10^{12\pm3}$  GeV (after inflation) Buchmuller, Yanagida, Plumacher, Ellis, Lola, Only survives if  $\Delta$ (B-L) is not zero Giudice et al, Fujii et al (otherwise is washed out at T<sub>ew</sub> by instantons) Main candidate: decay of lightest  $v_R$  (M~10<sup>12</sup> GeV) L non conserv. in  $v_{R}$  out-of-equilibrium decay: B-L excess survives at  $T_{ew}$  and gives the obs. B asymmetry. Quantitative studies confirm that the range of m<sub>i</sub> from v oscill's is compatible with BG via (thermal) LG In particular the bound  $m_i < 10^{-1} eV$ was derived for hierarchy Buchmuller, Di Bari, Plumacher; Can be relaxed for degenerate neutrinos Giudice et al; Pilaftsis et al; Sp\_fully compatible with oscill'n data!! Hambye et al

## Dark Matter

WMAP, SDSS, 2dFGRS....

Most of the Universe is not made up of atoms:  $\Omega_{tot} \sim 1$ ,  $\Omega_{b} \sim 0.045$ ,  $\Omega_{m} \sim 0.27$ Most is Dark Matter and Dark Energy

LHC

Most Dark Matter is Cold (non relativistic at freeze out) Significant Hot Dark matter is disfavoured Neutrinos are not much cosmo-relevant:  $\Omega_v < 0.015$ 

SUSY has excellent DM candidates: eg Neutralinos (--> LHC) Also Axions are still viable (introduced to solve strong CPV) (in a mass window around m ~10<sup>-4</sup> eV and  $f_a$  ~ 10<sup>11</sup> GeV but these values are simply a-posteriori)

Identification of Dark Matter is a task of enormous importance for particle physics and cosmology



LHC has good chances because it can reach any kind of WIMP:

WIMP: Weakly Interacting Massive Particle with m ~ 10<sup>1</sup>-10<sup>3</sup> GeV

For WIMP's in thermal equilibrium after inflation the density is:

$$\Omega_{\chi} h^2 \simeq const. \cdot \frac{T_0^3}{M_{\rm Pl}^3 \langle \sigma_A v \rangle} \simeq \frac{0.1 \ {\rm pb} \cdot c}{\langle \sigma_A v \rangle}$$

can work for typical weak cross-sections!!!

This "coincidence" is a good indication in favour of a WIMP explanation of Dark Matter

Strong competition from underground labs



Ð

A crucial question for the LHC

Is Dark Matter a WIMP?

LHC will probably tell yes or no to WIMPS



## Conceptual problems of the SM

Most clearly: • No quantum gravity (M<sub>Pl</sub> ~ 10<sup>19</sup> GeV)

 But a direct extrapolation of the SM leads directly to GUT's (M<sub>GUT</sub> ~ 10<sup>16</sup> GeV)

M<sub>GUT</sub> close to M<sub>PI</sub>



- suggests unification with gravity as in superstring theories
- poses the problem of the relation m<sub>w</sub> vs M<sub>GUT</sub>- M<sub>Pl</sub>

Can the SM be valid up to  $M_{GUT}$ -  $M_{PI}$ ?

The "big" hierarchy problem

Not only it looks very unlikely, but the new physics must be near the weak scale!

With new physics at  $\Lambda$  the SM is only an effective theory. After integration of the heavy d.o.f. :

 $\mathcal{L}_i$ : operator of dim i



Renorm.ble part

+  $o(1/\Lambda)\mathcal{L}_5$  +  $o(1/\Lambda^2)\mathcal{L}_6$  +...

Non renorm.ble part

In absence of special symmetries or selection rules, by dimensions  $c_i \mathcal{L}_i \sim o(\Lambda^{4-i}) \mathcal{L}_i$ 

 $\mathcal{L}_2$ : Boson masses  $\phi^2$ . In the SM the mass in the Higgs potential is unprotected:  $c_2 \sim o(\Lambda^2) = m_{W,H}$  should be  $o(\Lambda)$ !!  $\mathcal{L}_3$ : Fermion masses  $\overline{\psi}\psi$ . Protected by chiral symmetry and SU(2)xU(1):  $\Lambda$ ->mlog $\Lambda$  $\mathcal{L}_4$ : Renorm.ble interactions, e.g.  $\overline{\psi}\gamma^{\mu}\psi A_{\mu}$ 

 $\square_{4}$ : Non renorm.ble: suppressed by  $1/\Lambda^{i-4}$  e.g. $1/\Lambda^{2}\overline{\psi}\gamma^{\mu}\psi\overline{\psi}\gamma^{\mu}\psi$
# The "little hierarchy" problem



Another area where the SM is good, too good.....

With new physics at ~ TeV one would expect the SM suppression of FCNC and the CKM mechanism for CP violation to be sizably modified.

But this is not the case

an intriguing mystery and a major challenge for models of new physics



No clear signs of new physics in B decays (BaBar, Belle, Tevatron)



And now the LHCb experiment at the LHC has gone further in this direction

The CKM picture is confirmed as the main source of CPV in the quark sector

 $\ominus$  This poses strong constraints for models BSM

Adding effective operators to SM generally leads to very large  $\Lambda$  $M(B_{d}-\overline{B}_{d}) \sim \frac{(y_{t}V_{tb}*V_{td})^{2}}{16 \pi^{2} M_{w}^{2}} + c_{NP} \frac{1}{\Lambda^{2}}$ Isidori  $\sim 1 \xrightarrow{\text{tree/strong + generic flavour}} \Lambda \ge 2 \times 10^4 \text{ TeV [K]}$   $\sim 1/(16 \pi^2) \xrightarrow{\text{loop + generic flavour}} \Lambda \ge 2 \times 10^3 \text{ TeV [K]}$   $\sim (y_t V_{ti}^* V_{tj})^2 \xrightarrow{\text{tree/strong + MFV}} \Lambda \ge 5 \text{ TeV [K \& B]}$   $\sim (y_t V_{ti}^* V_{tj})^2/(16 \pi^2) \xrightarrow{\text{loop + MFV}} \Lambda \ge 0.5 \text{ TeV [K \& B]}$ 

But the hierarchy problem demands  $\Lambda$  in the few TeV range only assuming  $c_{NP} \sim (y_t V_{tb} * V_{td})^2$  (or anyway small) we get a bound on  $\Lambda$  in the TeV range

> eg in Minimal Flavour Violation (MFV) models D'Ambrosio, Giudice, Isidori, Strumia'02

# LHC and flavour physics Important results from LHCb



 $\oplus$ 

CMS & LHCb combined (presented at EPS'11 Grenoble)



 $\oplus$ 

A crucial question for the LHC

What damps the top loop  $\Lambda^2$  dependence?

• the s-top (SUSY)

some new fermion
t' (Little Higgs)
KK recurrences of the top (Extra dim.)

 nothing dumps it and we accept the ever increasing fine tuning



#### Solutions to the hierarchy problem

Supersymmetry: boson-fermion symm.

The most ambitious and widely accepted Simplest versions now marginal Plenty of viable alternatives

 Strong EWSB: Technicolor Strongly disfavoured by LEP. Coming back in new forms

> **Composite Higgs** Higgs as PG Boson, Little Higgs models.....

• Extra spacetime dim's that somehow "bring" M<sub>Pl</sub> down to o(1TeV) [large ED, warped ED, .....]. Holographic composite H Exciting. Many facets. Rich potentiality. No baseline model emerged so far

Ignore the problem: invoke the anthropic principle
Extreme, but not excluded by the data

#### The anthropic route

The scale of the cosmological constant is a big mystery.  $\Omega_{\Lambda} \sim 0.75 \qquad \rho_{\Lambda} \sim (2 \ 10^{-3} \ eV)^4 \sim (0.1 \text{ mm})^{-4}$ In Quantum Field Theory:  $\rho_{\Lambda} \sim (\Lambda_{\text{cutoff}})^4$  Similar to  $m_v$ !? If  $\Lambda_{\text{cutoff}} \sim M_{\text{Pl}} \qquad \rho_{\Lambda} \sim 10^{123} \rho_{\text{obs}}$ Exact SUSY would solve the problem:  $\rho_{\Lambda} = 0$ But SUSY is broken:  $\rho_{\Lambda} \sim (\Lambda_{\text{SUSY}})^4 \sim 10^{59} \rho_{\text{obs}}$ It is interesting that the correct order is  $(\rho_{\Lambda})^{1/4} \sim (\Lambda_{\text{FW}})^2/M_{\text{Pl}}$ 



"Quintessence" Λ as a vev of a field φ? Coupled to gauge singlet matter, eg v<sub>R</sub>, to solve magnitude and why now?

#### Is naturalness relevant? The multiverse alternative

Speculative physics reasons lead to doubts:

- The empirical value of the cosmological constant Λ poses a tremendous, unsolved naturalness problem yet the value of Λ is close to the Weinberg upper bound for galaxy formation
  - Possibly our Universe is just one of infinitely many continuously created from the vacuum by quantum fluctuations
  - Different physics in different Universes according to the multitude of string theory solutions (~10<sup>500</sup>)

Perhaps we live in a very unlikely Universe but one that allows our existence Given the stubborn refuse of the SM to step aside, and the terrible unexplained naturalness problem of the cosmological constant, many people have turned to the anthropic philosophy also for the SM

I find applying the anthropic principle to the SM hierarchy problem still completely unmotivated

After all, we can find plenty of models that reduce the fine tuning from 10<sup>14</sup> to 10<sup>2</sup>: so why make our Universe so terribly unlikely?

The case of the cosmological constant is a lot different: the context is not as fully specified as the for the SM (quantum gravity, string cosmology, branes in extra dims., wormholes thru different Universes....)



An example of anthropic picture

An enlarged SM (to include RH  $\nu$ 's and no new physics) remains as an open but enormously fine tuned option

A light Higgs

SO(10) non SUSY GUT

SO(10) breaking down to  $SU(4)xSU(2)_LxSU(2)_R$  at an intermediate scale (10<sup>11-12</sup>)

Majorana neutrinos and see-saw (->  $0\nu\beta\beta$ )

Axions as dark matter

Baryogenesis thru leptogenesis

Æ

But:  $(g-2)_{\mu}$  and other present hints of deviations from SM should disappear or be explained away

# Some amount of new physics could bring EW precision tests better into focus



## Muon g-2

 $a_{\mu}$  is a plausible location for a new physics signal!!

χ

eg could be light SUSY (now tension with LHC)

```
a_{\mu}^{exp} - a_{\mu}^{SM} = (28.7 \pm 8.0) \times 10^{-10}
```

- 3.6 "standard deviations" ( $e^+e^-$ )
- 2.4 "standard deviations" (τ)

$$\delta a_{\mu} = 13 \cdot 10^{-10} \left(\frac{100 GeV}{M_{SUSY}}\right)^2 tg\beta$$



Some NP hints from accelerator experiments

A <sup>b</sup> <sub>FB</sub> LEP		~	-3σ	
(g-2) <sub>µ</sub> Broc	okhaven	~	-3σ	
tt <sup>bar</sup> FB asymmetry	Tevatron (m	ostly CDF)	~3 o at large M	tt
Dimuon charge asyr	nmetry	DO	~ <b>3.9</b> σ	
Wjj excess at M <sub>jj</sub> ~ 14 only candidate to o	4 GeV pen prod. of NP	CDF not con	<b>~3.2</b> σ firmed by D0, LHC	
$B_s \rightarrow J/\psi\phi$	Te	evatron, LHC	Cb ~went away	
$B \rightarrow \tau v$	В	aBar, Belle	~2.5o	
CPV in D->ππ, KK Lŀ	ICb All o	f them cou	ld still go away!	

#### A non-LHC very important result

MEG new limit on Br( $\mu \rightarrow e \gamma$ ) < 2.4 10<sup>-12</sup>



No neutron electric dipole moment

d<sub>n</sub> violates P and T

$$\vec{d}_n = d\vec{\sigma}$$
  $\vec{m}_n = \mu\vec{\sigma}$ 

$$H \sim -(\vec{d}_n \cdot \vec{E} + \vec{m}_n \cdot \vec{B}) = -(d\vec{E} + \mu \vec{B}) \cdot \vec{\sigma}$$

E and B have opposite behaviour under P and T CPT is conserved, so T violation implies CP violation

Present limit on d<sub>n</sub> from Grenoble

 $|\mathbf{d}_n| < 3 \ 10^{-26} \ \mathrm{e} \ \mathrm{cm} \ (90\% \mathrm{cl})$ 



A striking result of the 2011 LHC run ( > 1 fb<sup>-1</sup>) is that the new physics is pushed further away

But only ~ 20-25% of the 2011 statistics has been analysed

Examples:

sequential W':  $m_{W'} > 2.3$  TeV sequential Z':  $m_{Z'} > 1.9$  TeV axi-gluon: 2.5-3.2 TeV gluino:  $m_g > \sim 0.5 - 1$  TeV

Many generic signatures searched. Not a single significant hint of new physics found



### **Di-lepton Channel**



Sequential SM: m(Z') > 1.9 TeV at 95% C.L. RS graviton ( $k/M_{Pl} = 0.1$ ): m(G) > 1.8 TeV at 95% C.L.



# **Di-photon Channel**

┢

## RS graviton (k/MPI = 0.1): m(G) > 1.7 TeV at 95% C.L.



 $W' \rightarrow V$ 

# Sequential SM: m(W') > 2.3 TeV at 95% C.L.



Events

 $\oplus$ 

# Dijet

Model	95% CL Limits (TeV)		
ATL-CONF-2011-095	Expected	Observed	
Excited Quark $q^*$	2.77	2.91	
Axigluon	3.02	3.21	
Color Octet Scalar	1.71	1.91	

Model	Excluded Mass (TeV)		
CMS arXiv.1107.4771	Observed	Expected	
String Resonances	4.00	3.90	
E <sub>6</sub> Diquarks	3.52	3.28	
Excited Quarks	2.49	2.68	
Axigluons/Colorons	2.47	2.66	
W' Bosons	1.51	1.40	



### SUSY: boson fermion symmetry

The hierarchy problem:  $\delta m_{h|top}^2 = -\frac{3G_F}{2\sqrt{2}\pi^2}m_t^2\Lambda^2 \sim -(0.2\Lambda)^2$ 

In broken SUSY  $\Lambda^2$  is replaced by  $(m_{stop}{}^2\text{-}m_t{}^2)\text{log}\Lambda$ 

 $m_H$ >115.5 GeV,  $m_{\chi+}$ >100 GeV, EW precision tests, success of CKM, absence of FCNC, all together, impose sizable Fine Tuning (FT) particularly on minimal realizations (MSSM, CMSSM...).

Yet SUSY is a completely specified, consistent, computable model, perturbative up to  $M_{Pl}$  quantitatively in agreement with coupling unification (GUT's) (unique among NP models) and has a good DM candidate: the neutralino (actually more than one).

Remains the reference model for NP

# Beyond the SM SUSY is unique in providing a perturbative theory up to the GUT/Planck scale



Other BSM models (little Higgs, composite Higgs, Higgsless....) all become strongly interacting and non perturbative at a multi-TeV scale

#### SUPERGAUGE TRANSFORMATIONS IN FOUR DIMENSIONS

J. WESS Karlsruhe University

B. ZUMINO

CERN, Geneva

Received 5 October 1973

Abstract: Supergauge transformations are defined in four space-time dimensions. Their commutators are shown to generate  $\gamma_5$  transformations and conformal transformations. Various kinds of multiplets are described and examples of their combinations to new representations are given. The relevance of supergauge transformations for Lagrangian field theory is explained. Finally, the abstract group theoretic structure is discussed.

Volume 49B, number 1

PHYSICS LETTERS

18 March 1974

#### A LAGRANGIAN MODEL INVARIANT UNDER SUPERGAUGE TRANSFORMATIONS

J. WESS

Karlsruhe University, Germany

and

B. ZUMINO CERN, Geneva, Switzerland

Received 4 January 1974

We study, in the one-loop approximation, a Lagrangian model invariant under supergauge transformations. The model involves a scalar, a pseudoscalar and a spinor field. Supergauge invariance gives rise to relations among the masses and the coupling of these fields and implies the existence of a conserved current. The renormalization procedure is discussed and the relations among masses and couplings are shown to be preserved by renormalization.



The Wess- Zumino model is the basis for the MSSM, central in the LHC programme Julius Wess: a teacher

Most theorists have learnt SUSY from this book Julius Wess Jonathan Bagger

#### Supersymmetry and Supergravity

SECOND EDITION REVISED AND EXPANDED

PRINCETON SERIES IN PHYSICS



The general MSSM has > 100 parameters

Simplified versions with a drastic reduction of parameters are used for practical reasons, e.g.

CMSSM, mSUGRA : universal gaugino and scalar soft terms at GUT scale  $m_{1/2}$ ,  $m_0$ ,  $A_0$ ,  $tg\beta$ ,  $sign(\mu)$ 

NUHM1,2: different than  $m_0$  masses for  $H_u$ ,  $H_d$  (1 or 2 masses)

It is only these oversimplified models that are now cornered



### Jets + missing E<sub>T</sub>

#### CMSSM (degenerate s-quarks)



 $\oplus$ 



Impact of  $m_H \sim 125$  GeV on SUSY models

Simplest models with gauge mediation are disfavoured (predict  $m_H$  too light)

Djouadi et al; Draper et al, '11

some versions, eg gauge mediation with extra vector like matter, or with Higgs-messenger mixing do work

Endo et al '11, Evans et al '12

Anomaly mediation is also generically in trouble

Gravity mediation is better but CMSSM, mSUGRA, NUHM1,2 need squarks heavy, A<sub>t</sub> large and lead to tension with g-2 (that wants light SUSY) and b->s $\gamma$ 

Akura et al; Baer et al; Battaglia et al; Buchmuller et al, Kadastik et al; Strege et al; '11



#### maximal top mixing is required

+





Baer et al '11

 $\oplus$ 



Input data for existing fits of CMSSM, NUHM1 ..... include

- The EW precision tests
- Muon g-2

- e.g. MASTERCODE Buchmuller et al
- Flavour precision observables
- Dark Matter
- Higgs mass constraints and LHC

Buchmuller et al '11

Pre LHC '11 fit

 $\pm$ 





With 2011 LHC data heavier scalars

Tension g-2 vs m<sub>H</sub> ~ 125 GeV





with g-2  $m_H \sim 119$  GeV without g-2  $m_H \sim 125$  GeV

g-2 indicates light SUSY!!

 $\oplus$
With new data ever increasing fine tuning

One must go to SUSY beyond the CMSSM, mSUGRA, NUHM1,2

There is still room for more sophisticated versions

- Heavy first 2 generations
- NMSSM
- $\lambda$  SUSY
- Split SUSY
- Large scale SUSY
- • •

Beyond the CMSSM, mSugra, NUHM1,2

## Heavy 1st, 2nd generations





## For example, may be gluinos decay into 3-gen squarks



An extra singlet Higgs

In a promising class of models a singlet Higgs S is added and the  $\mu$  term arises from the S VEV (the  $\mu$  problem is solved)

 $\lambda SH_uH_d$ 

Mixing with S can bring the light Higgs mass down at tree level (no need of large loop corrections) Depending on the value of  $\lambda$ :







¢



It is not excluded that at 125 GeV you see the heaviest of the two and the lightest escaped detection at LEP Ellwanger '11

Mixing with S makes h light already at tree level No need of loops Fine tuning can be very small 3.0  $m_{h}^{2} < 0$ 2.5 (g) 2.0  $m_h$ 1.5 1.0

0.8

1.0

1.2

14

λ

1.6

18

20

In MSSM it is not possible to obtain an enhanced  $\gamma\gamma$  signal for m<sub>H</sub> ~ 125 GeV, while it is possible eg in NMSSM or  $\lambda$  SUSY Arvanitaki et al, Hall et al '11



Drawback of  $\lambda$  SUSY: relation with GUT's & coupling unification is generically lost

(+)

## Conclusion

## LHC scenarios

Catastrophic: No Higgs, no new physics Can only occur if the LHC is not enough to fully probe the EW scale: unitarity violations impose one or the other (eg new vector bosons) or both

The Higgs comes closer: yes or no to the SM Higgs in 2012 Hints (to be confirmed) of  $m_{\rm H} \sim 125~{\rm geV}$ 

Theorist projection: non standard Higgs and new physics

A lot of model building in this direction

No new Physics so far: but LHC is just at beginning

Pure SM: A light scalar Higgs, no new physics at the LHC If so, nature does not at all abhor fine tuning This is the anthropic paradigm that experiment must try to falsify

