

# CRAYS: a photometric calibration of TA FD camera

Sept. 14<sup>th</sup> 2011, 8<sup>th</sup> AFWS @ KIT

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# Determination of FD Energy

## Systematic uncertainty

Source	$\Delta E/E$
Fluorescence yield	11%
Detector	10%
Atmosphere	11%
Reconstruction	10%
Total	21%

2. Atmospheric Correction

$E_i$

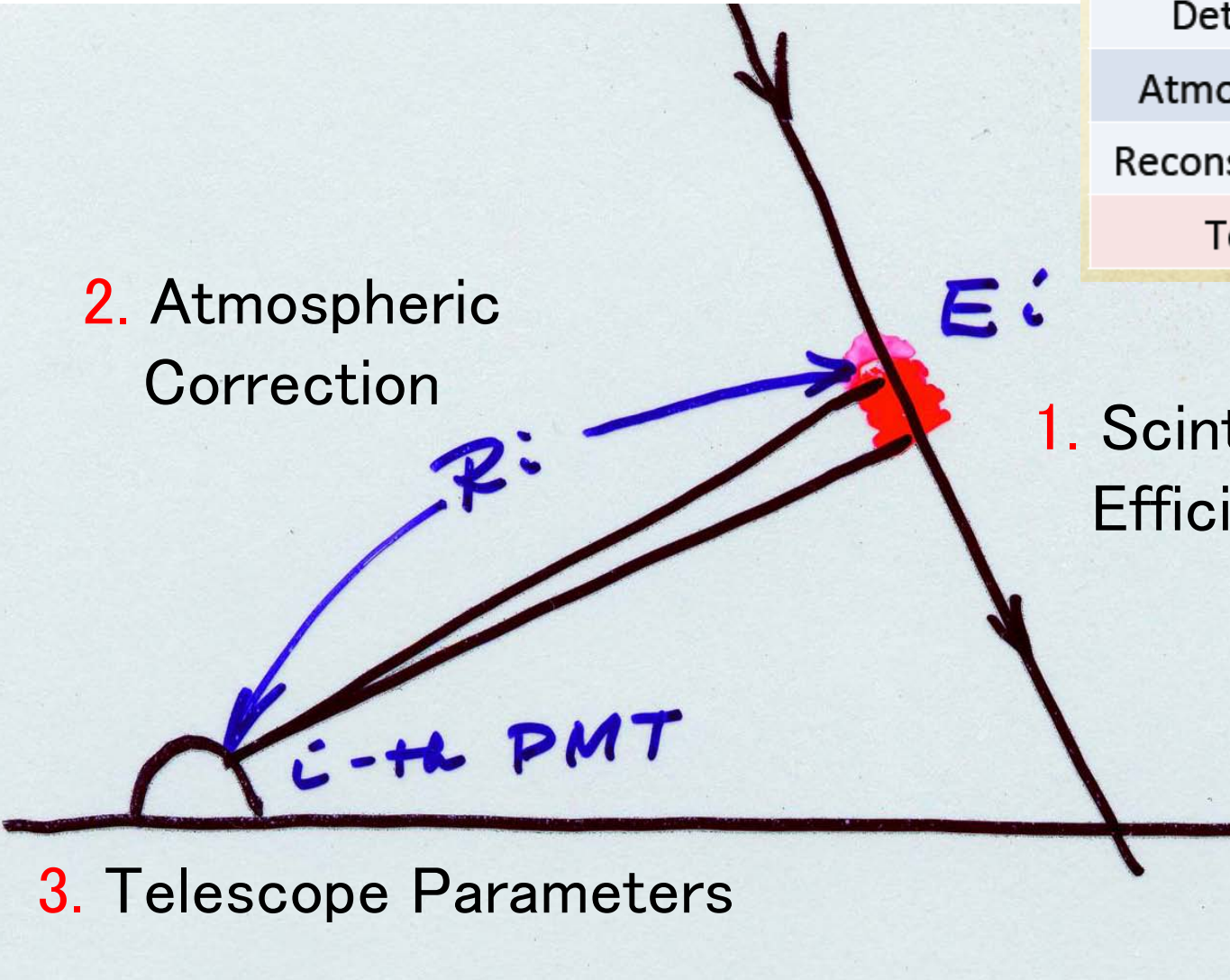
1. Scintillation Efficiency

$R_i$

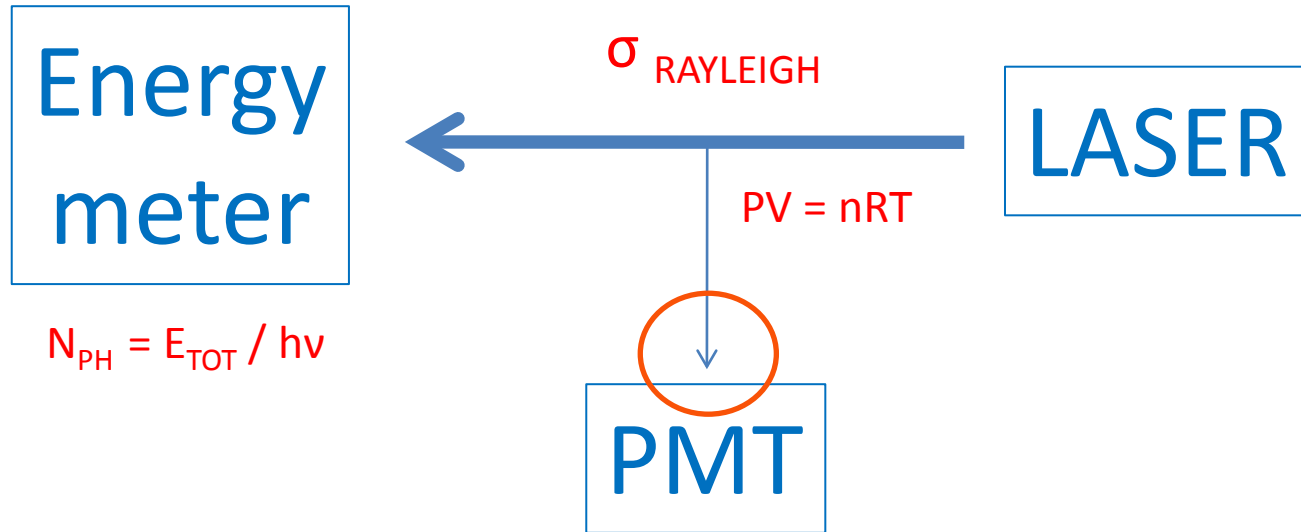
$L \rightarrow$  the PMT

3. Telescope Parameters

4. Loss of HE  $\mu$  &  $\nu$



# How to Calibrate?



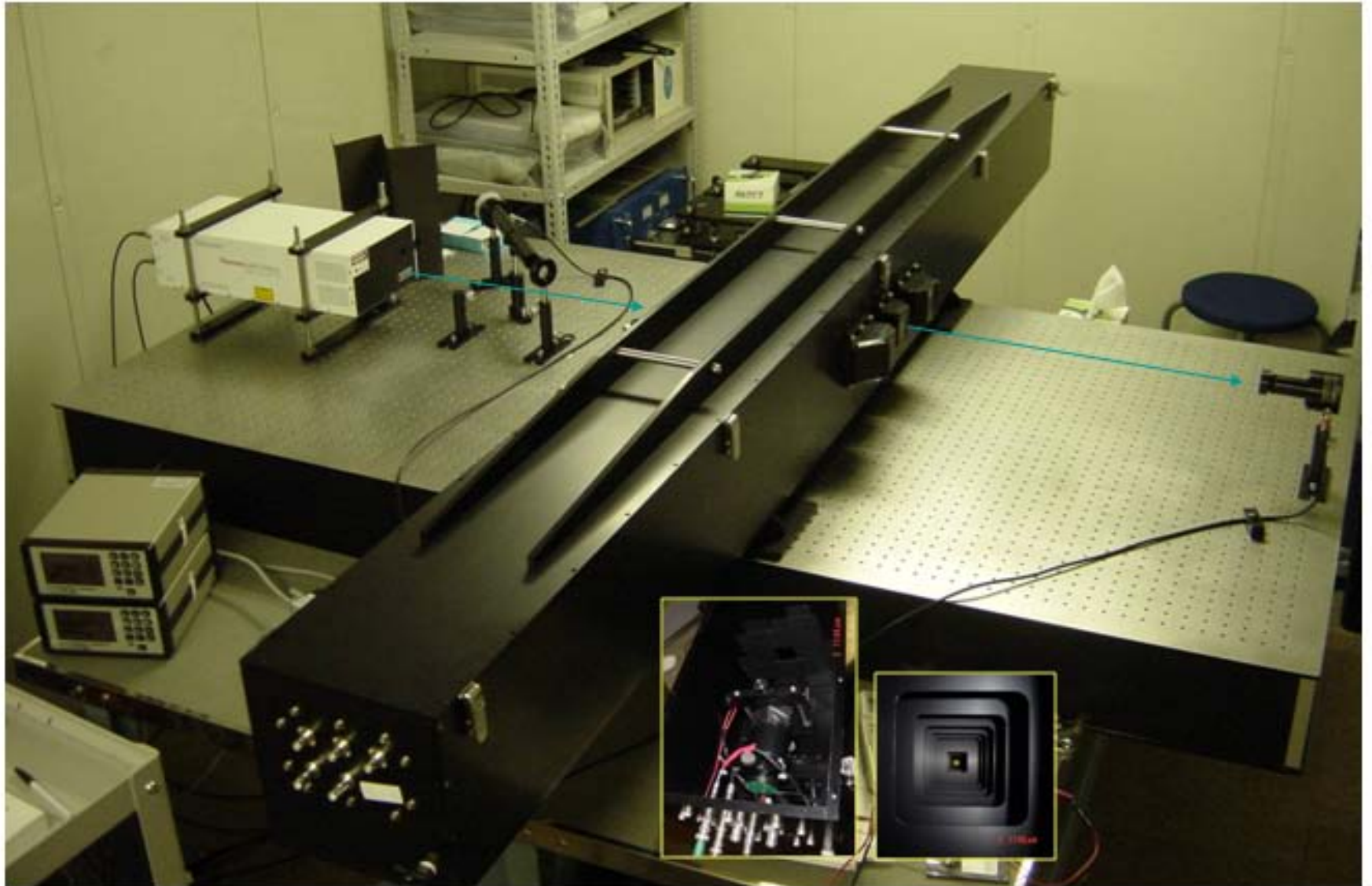
## Absolutely calibrated light source

- Pulsed ( $\sim 4\text{ns}$ )
- UV ( $\sim 337\text{nm}$ )
- Large # of photons ( $\sim 100 \text{ photons/cm}^2$ )
- Linear with pressure ( $\sim 0 \text{ photons/cm}^2$ )

# CRAYS : Calibration using RAYleigh Scattering

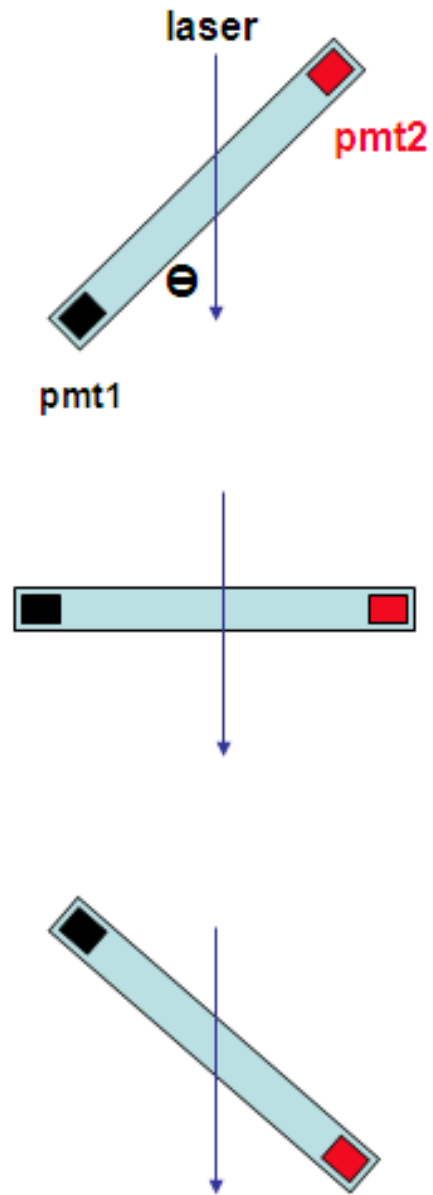
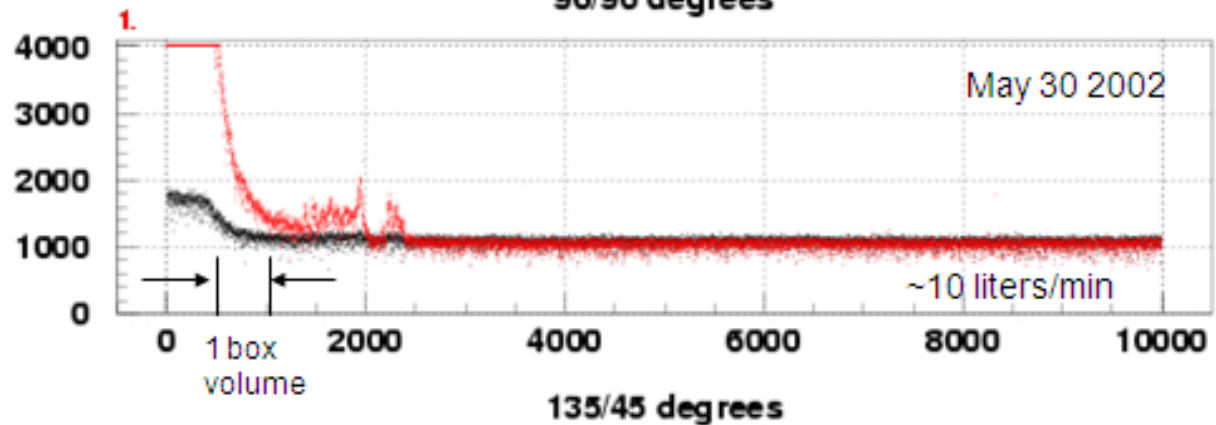
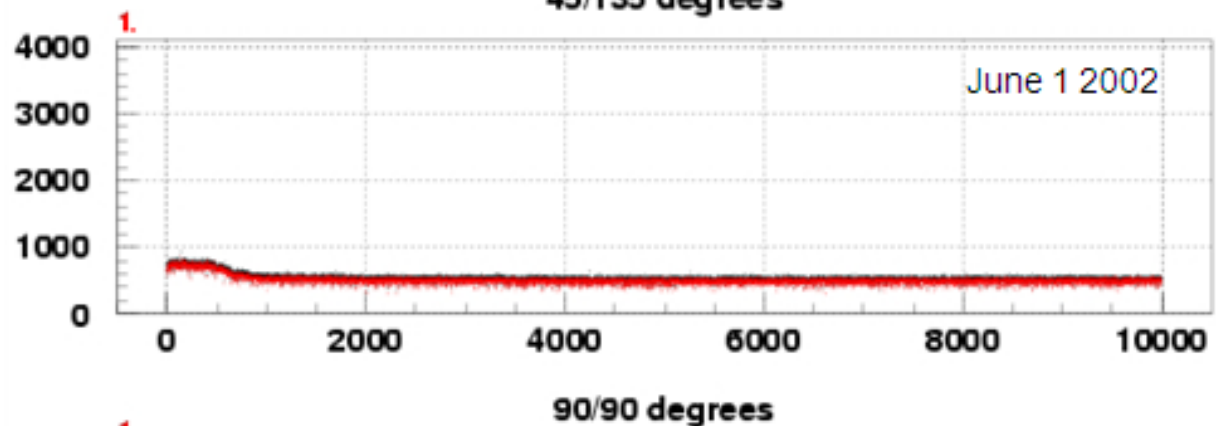
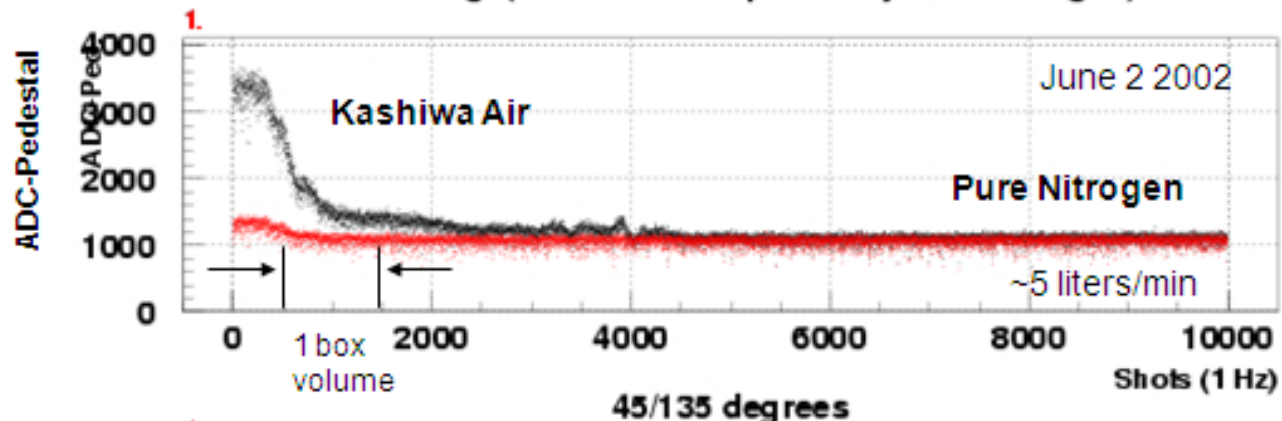
- Original idea by discussion with L.Wiencke (2002)
- H.Naus & W.Ubachs: Optics Lett.25 (2000)  
Rayleigh scattering's  $\sigma$  measured with 1% accuracy
- CRAYS-I by L.Wiecke (2002)
- CRAYS-II by N.Sakurai (2004)
- IceCube PMT calib. in 1-ph mode by S.Yoshida (2007)
- TA PMT calib. in full intensity mode by S.Kawana (2008)

# CRAYS-I

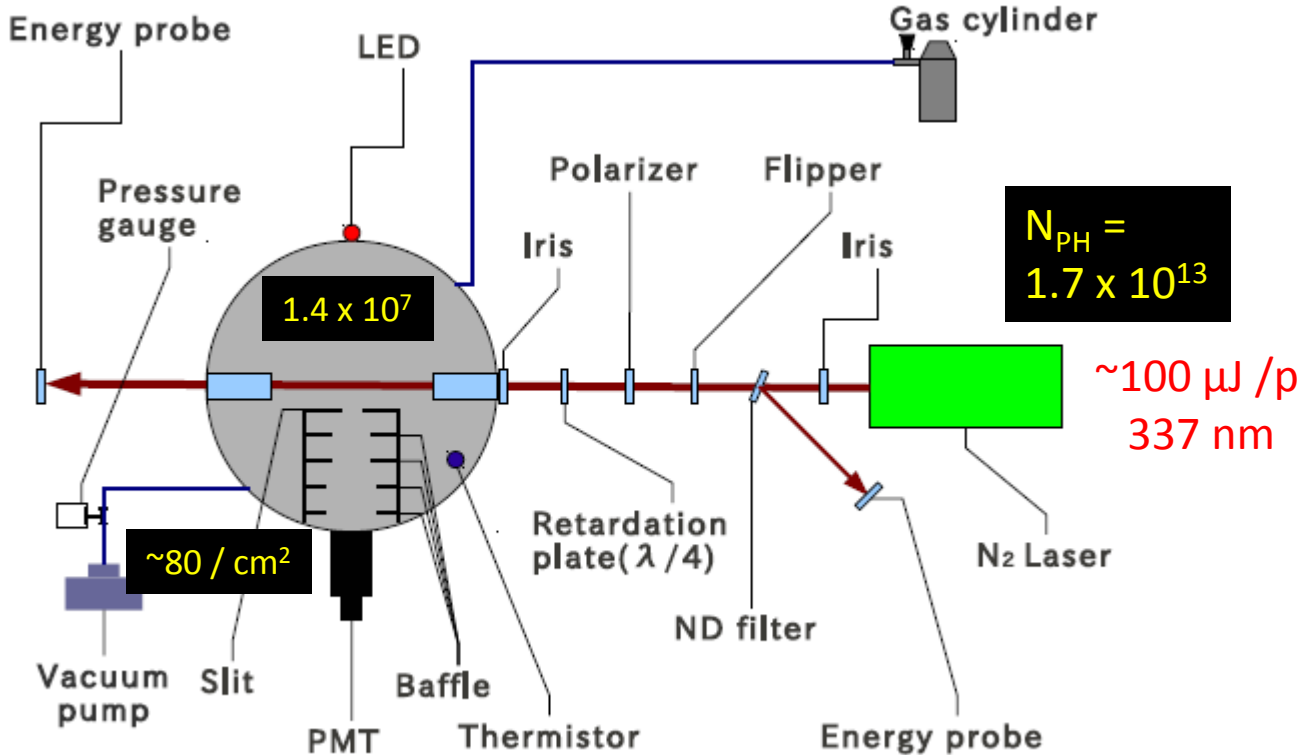
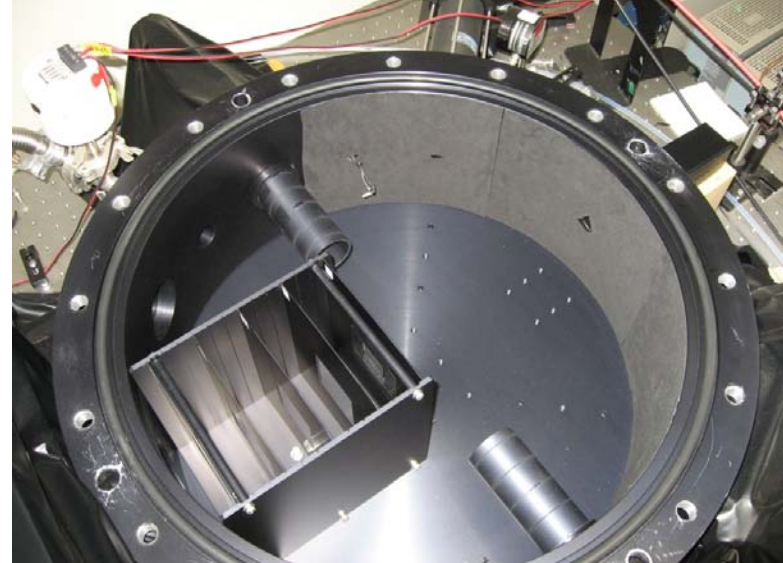


# CRAYS-I

Scattering (Kashiwa Air replaced by Pure Nitrogen)

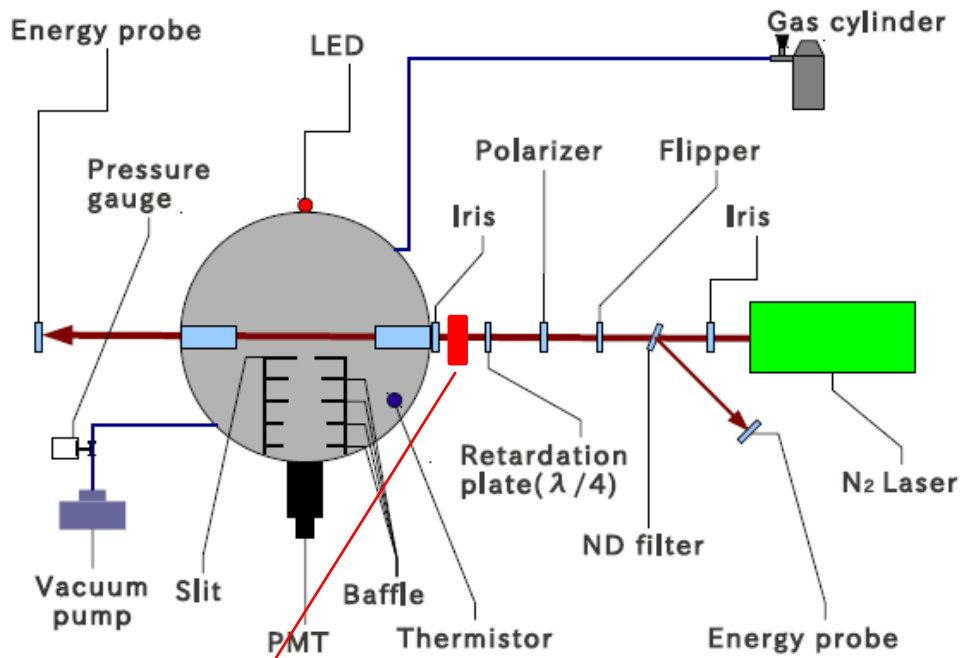


# CRAYS-II



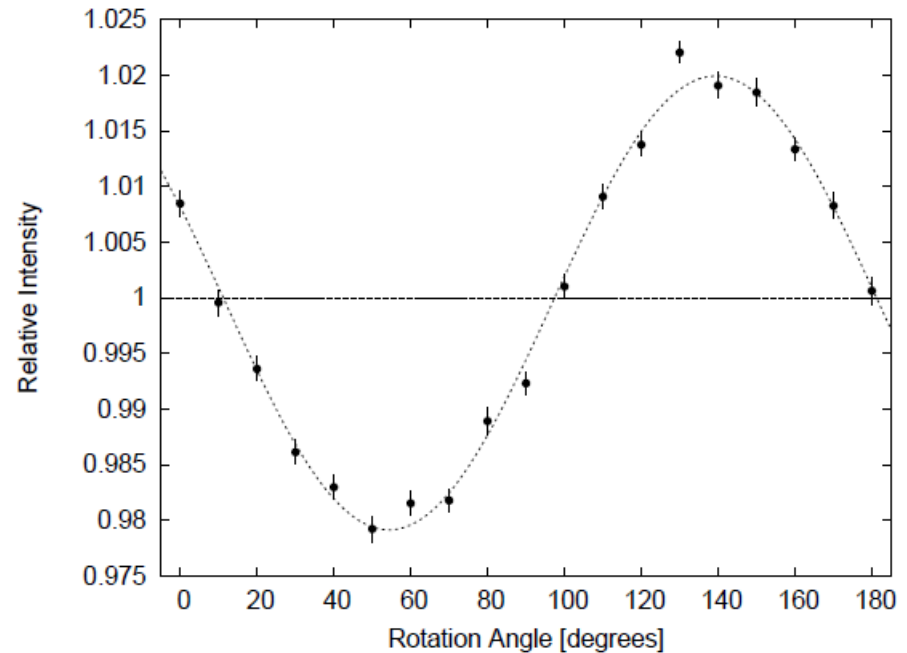
CaF<sub>2</sub> with AR coat used for all windows

# Check-1: polarization



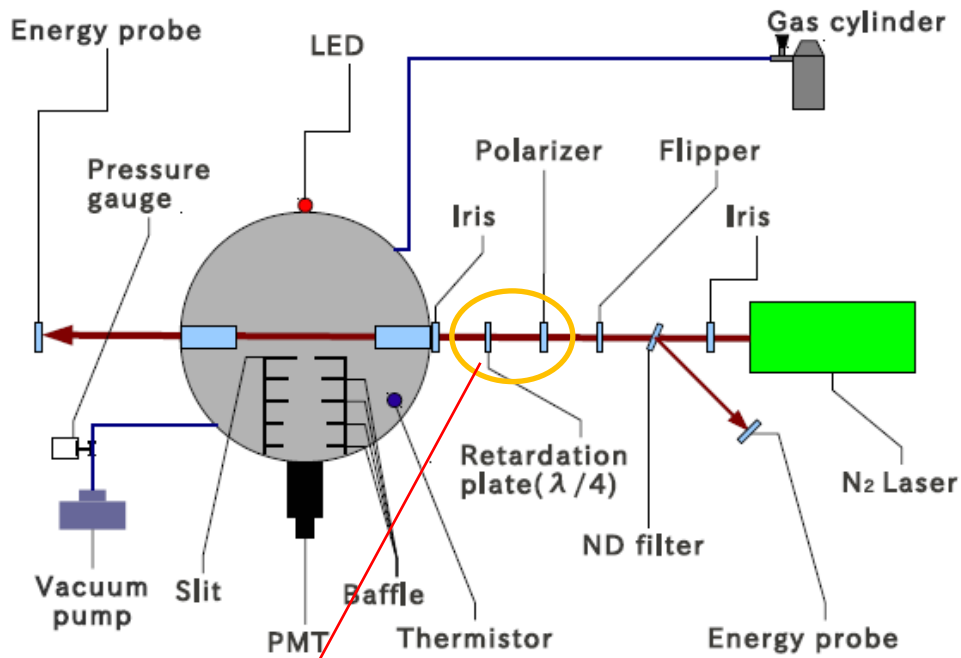
1. Polarizer +  $\lambda/4$  produces circular pol..
2. Insert another polarizer after  $\lambda/4$
3. Rotate pol. angle & measure laser energy

2% vertical pol.  
➤ No effect  
on X-section



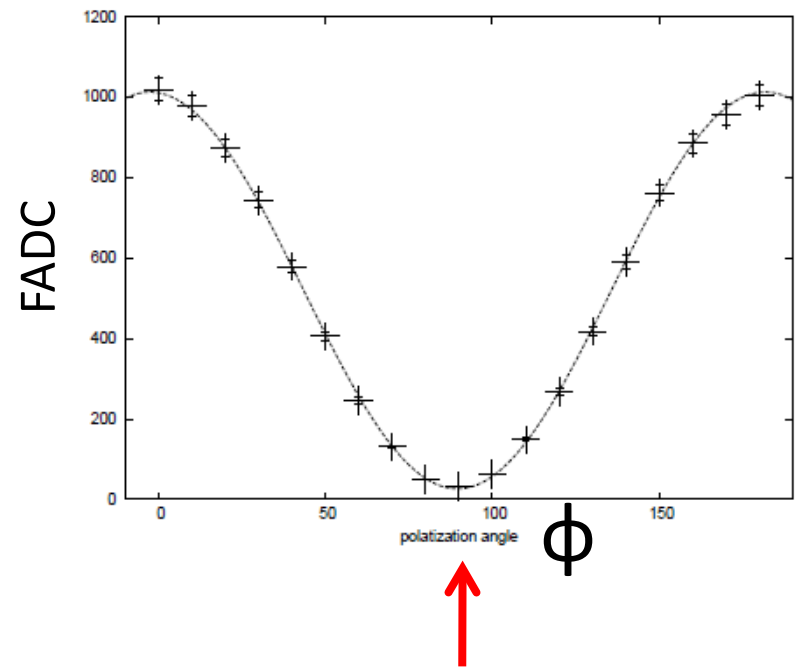


# Check-2: Rayleigh $\phi$ dep?



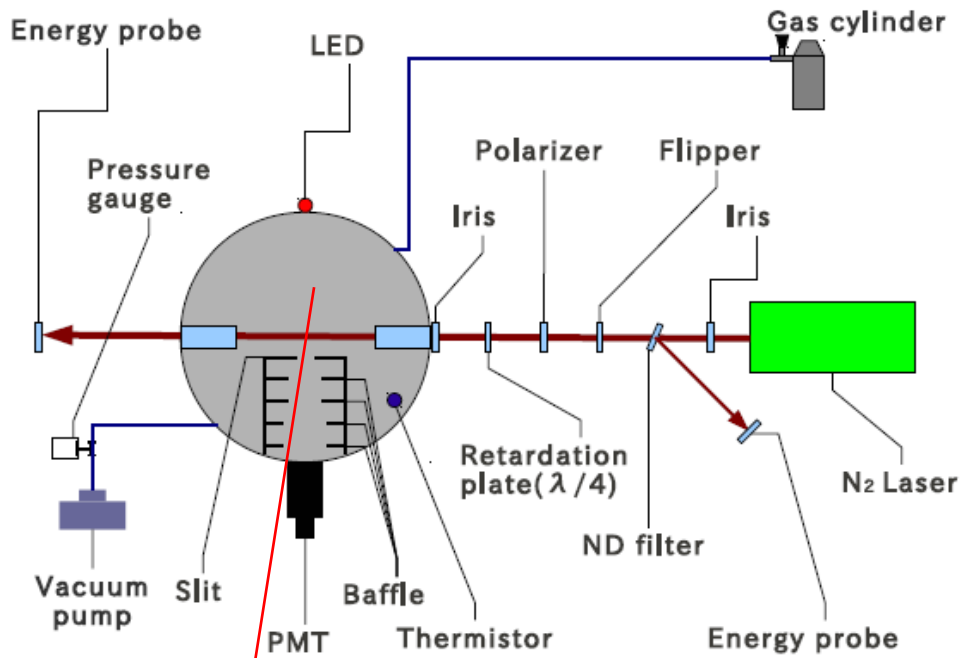
1. Remove  $\lambda/4$  and make full linear pol..
2. Rotate pol. angle & measure PMT output.

$$\propto \cos(2\phi)$$



Remaining  $\sim 3\%$ : BG +  
 $N_2$  depolarization effect

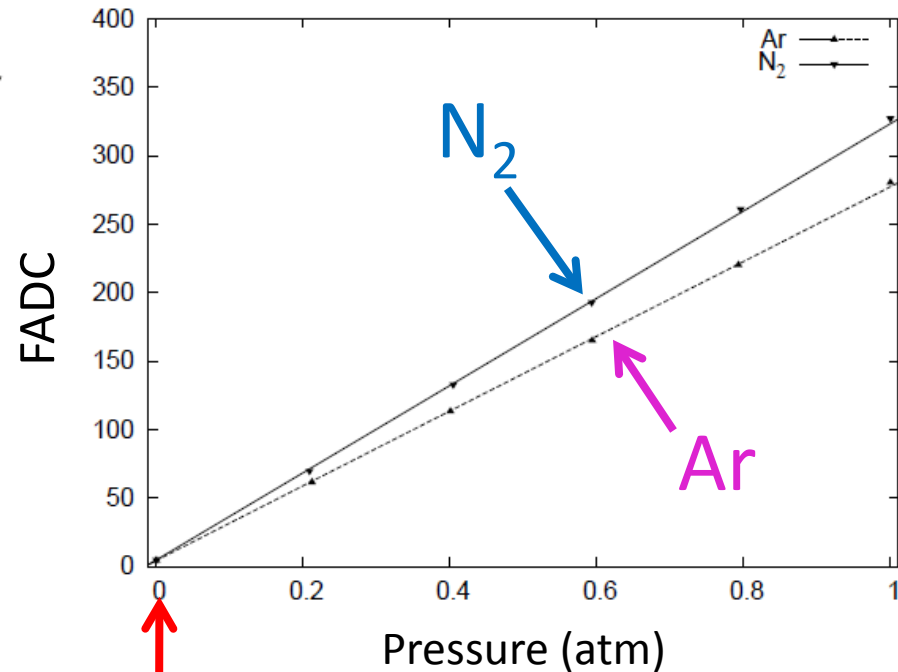
# Check-3: Pressure curve & Ar/N<sub>2</sub> ratio



1. Fill N<sub>2</sub> and change P, measure PMT output
2. Repeat the same with Ar

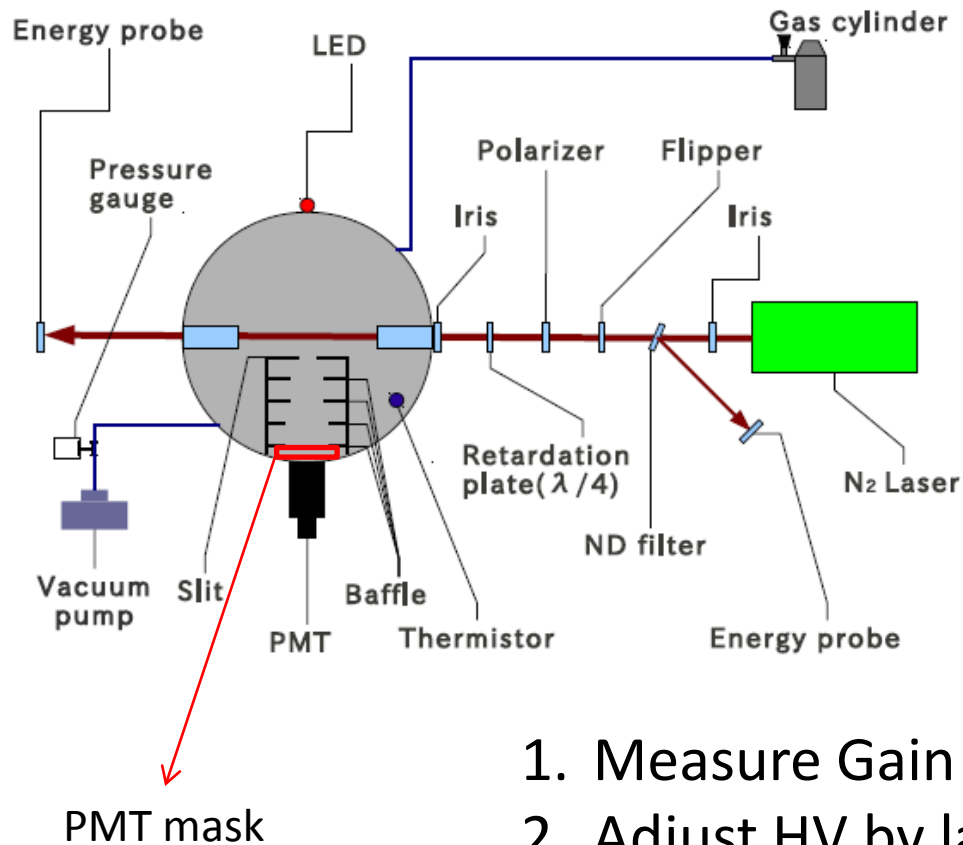
Rayleigh scattering of N<sub>2</sub> laser on N<sub>2</sub> gas is OK.

$$\text{Ar/N}_2 = 0.857 \pm 0.007$$
$$\text{theory} = 0.859$$



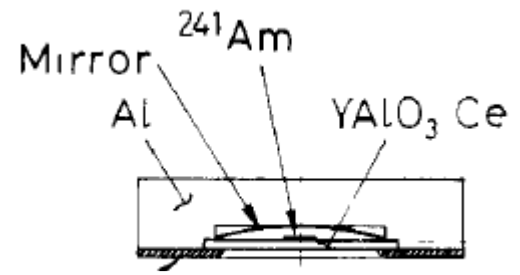
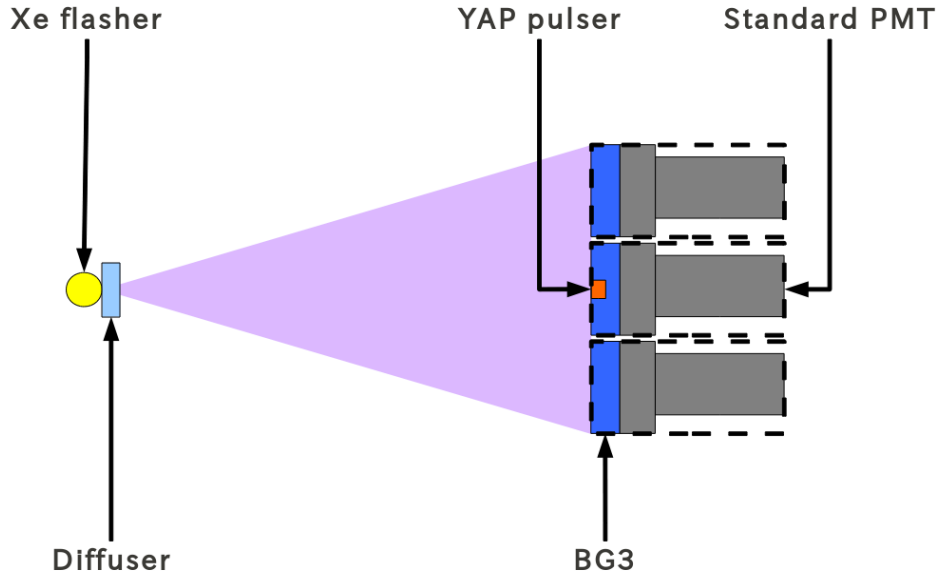
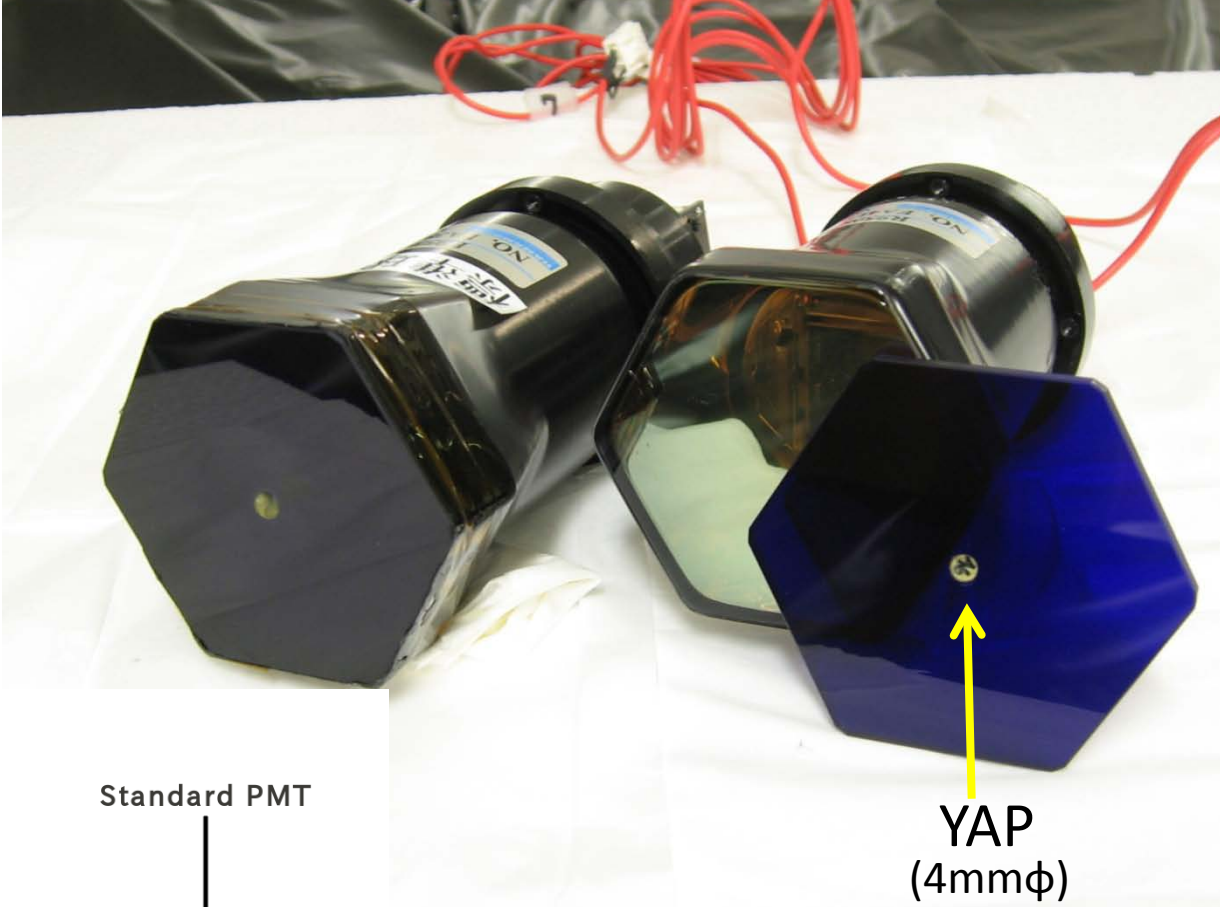
BG < 2% at 0 atm

# Calibration Procedure for 75 PMTs with YAP



1. Measure Gain sensitivity %/V by UV-LED
2. Adjust HV by laser for targeted calib. value  
~ 2.3 [photons /  $\Sigma$ FADC counts]
3. Measure with 20mm $\phi$ , 36mm $\phi$  mask and  
No mask
4. Measure YAP for future PMT gain tracing

# Standard PMT, YAP & Xenon flasher



# Signal integration

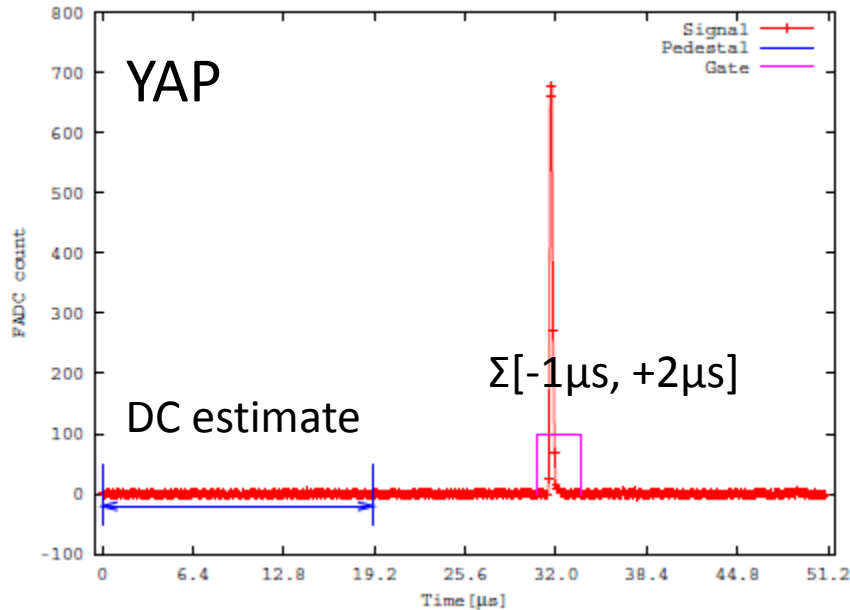
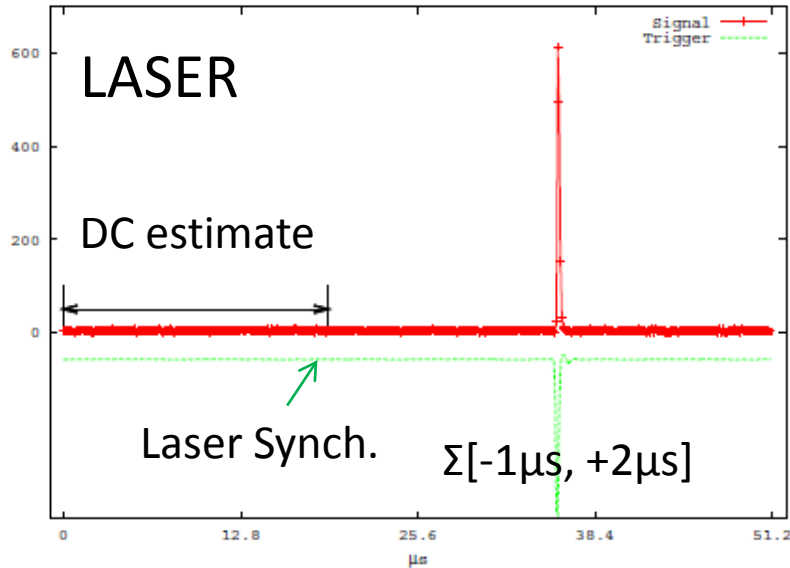
## Triple subtraction

1. Pedestal DC
2. Flipper-IN (no laser light)
3. Vacuum (no Rayleigh scattering)

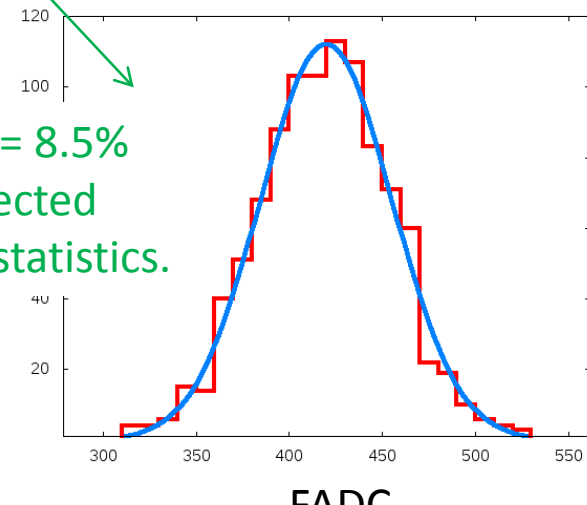
Flipper-OUT:  
10 x 100 ev

Flipper-IN:  
10 X 100 ev

$$\Sigma(\text{FADC} - \text{dc}) - \Sigma(\text{FADC} - \text{dc})$$



$\sigma/\text{peak} = 8.5\%$   
7% expected  
by p.e. statistics.



# Target (N<sub>2</sub>) Density

- $n/V$  by state equation;  $PV = nRT$
- van der Waals correction negligible
- Chamber  $P =$   
gauge- $P +$  abs.- $P$  by mercury gauge.
- Gas temp. measured.

# # of expected photons by ray tracing

## Rayleigh scattering cross section †

- Theoretical formula:

$$\sigma_R(\nu) = \frac{24\pi^3\nu^4}{N^2} \left( \frac{n(\nu)^2 - 1}{n(\nu)^2 + 2} \right)^2 F_k(\nu)$$

$$F_k(\nu) = 1.034 + 3.17 \times 10^{-12}\nu^2$$

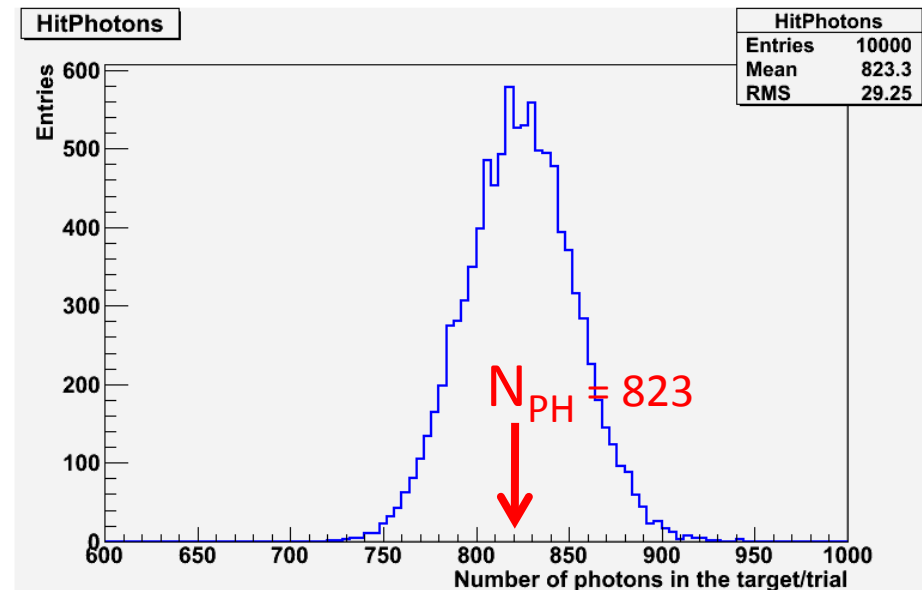
$$10^8[n(\nu) - 1] = 5989.242 + \frac{3.3632663 \times 10^{14}}{1.44 \times 10^{10} - \nu^2} \quad (256 \leq \lambda \leq 468[\text{nm}])$$



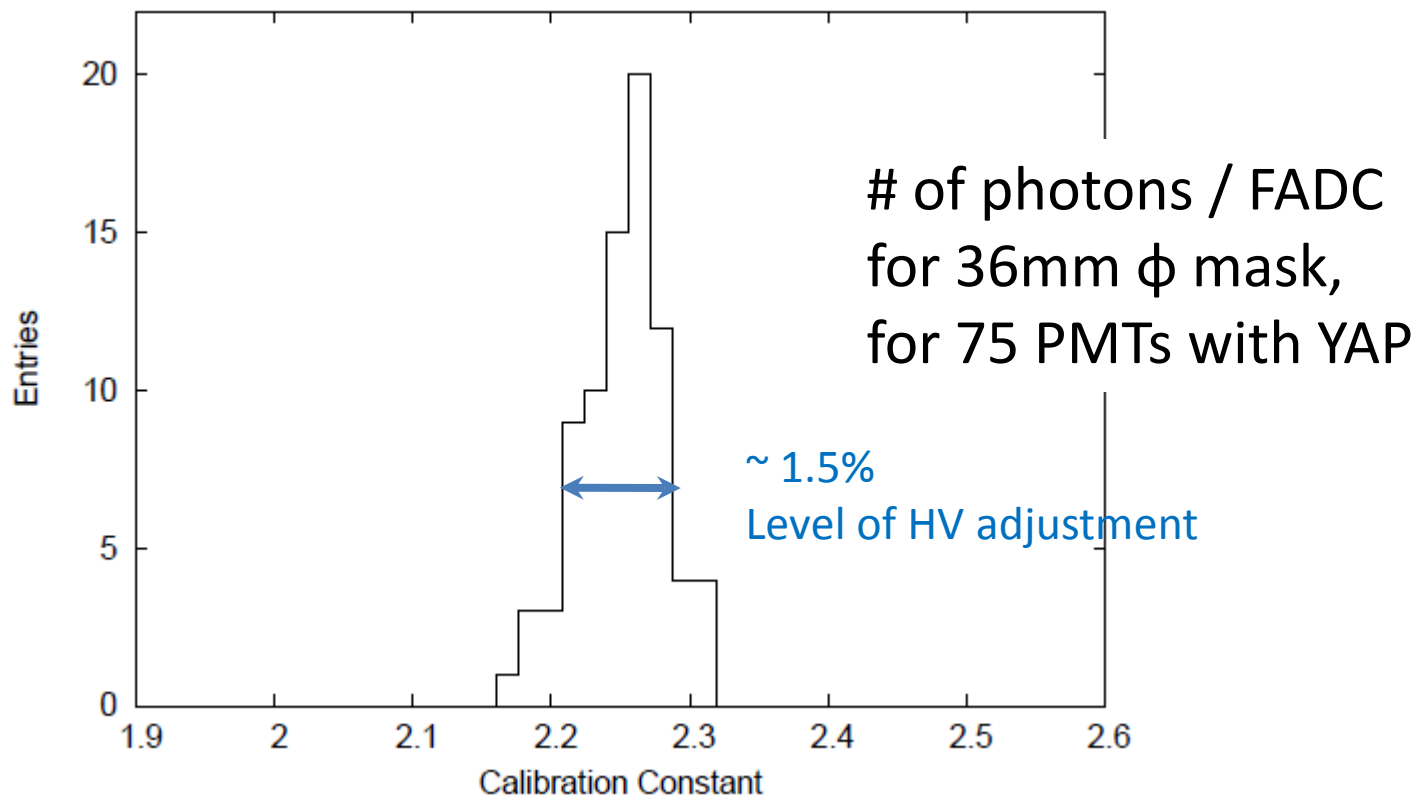
$$\frac{d\sigma}{d\Omega} = \frac{3}{16\pi} (1 + \cos^2 \theta) \times 3.50 \times 10^{-26} \text{ cm}^2$$

- 200 nJ laser pulse
- N<sub>2</sub> gas at 25°C
- 1000 hPa
- Given geometry (36mm  $\phi$  PMT mask)
- Wall reflection :  
2.3% random (or mirror) ref.

$$N_{\text{BG}} < 1$$



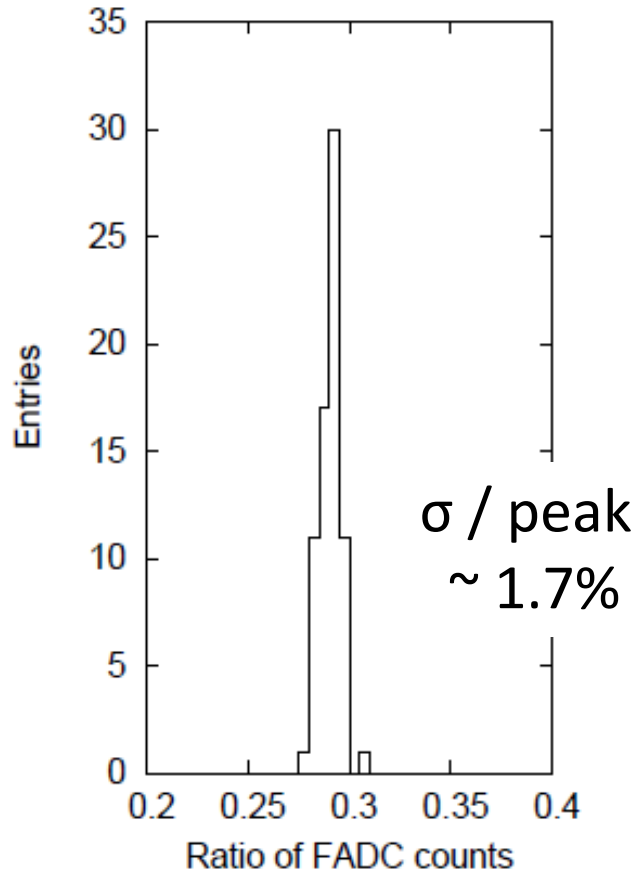
# Determined Calibration Constant



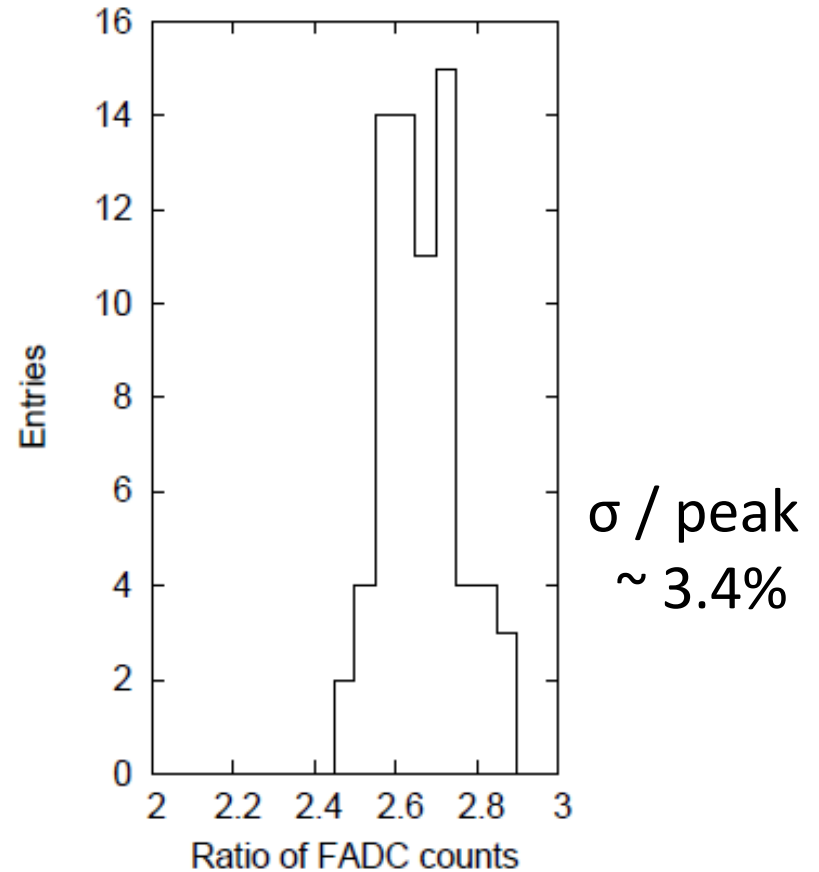
2.25 Photons / FADC cnt  $\sim$  PMT Gain of  $6 \times 10^4$



# FADC with different PMT mask



20mm $\phi$  mask / 36mm $\phi$  mask



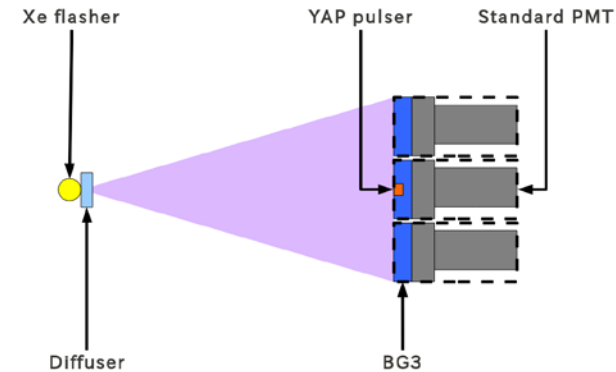
no mask / 36mm $\phi$  mask

# Scale Uncertainty of CRAYS

	Error
Rayleigh scattering cross-section	1.0%
Molecular density (temperature and pressure)	2.0%
Measurement of laser energy	5.0%
Polarization of the laser beam	1.0%
Geometric aperture calculation	3.0%
Signal integration	2.0%
Background and noise subtraction	2.0%
Effect of geomagnetism	1.0%
Total (quadratic sum of above)	7.0%

# Stability (Kashiwa > Utah)

- CRAYS-calibrated PMTs transported to Utah.
- 2PMTs / camera installed with same HV.
- PMT @ center = standard PMT.

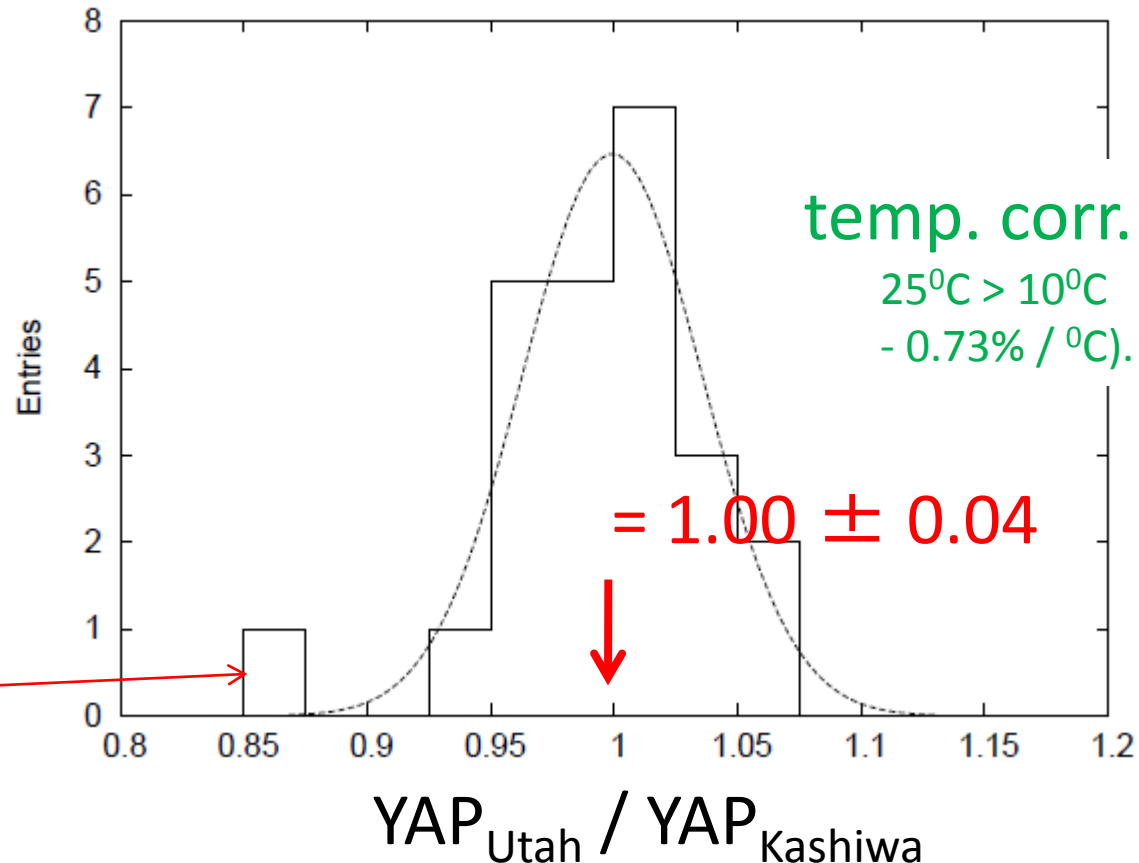


## Kashiwa > Utah

1. HVPS
2. FADC and cable
3. Temperature
4. Geomagnetism etc. are different.

+ PMT gain + YAP light

one PMT with 0.85  
(YAP problem)



# Summary

- TA FD camera: CRAYS calibration with YAP gain monitoring
- Abs. Scale Uncertainty  $\sim 0.07$
- Calib. transport from Kashiwa to Utah  $= 1.00 \pm 0.04$

not discussed today:

- Extension from standard PMT to other 255 PMTs by diffused xenon flasher
- Long term gain monitoring in situ (PMT aging?)
- Cross calibration with Electron Light Source (ELS)

$$\text{AF Yield} \times \overset{\text{good}}{\text{CRAYS}} \times \overset{\text{this}}{\text{FF}} \overset{1.000?}{=} \text{ELS (40MeV electron linac).}$$

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1. Scintillation Efficiency

$R_i$

$L \rightarrow$  the PMT

3. Telescope Parameters

4. Loss of HE  $\mu$  &  $\nu$

